

# Superaligned Gamow-Teller Decay of Doubly Magic $^{100}\text{Sn}$

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## Abstract

abstract

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## 1 I. SUMMARY

2 The shell structure of atomic nuclei is associated with magic numbers and originates from  
3 the nearly independent motion of neutrons and protons in a mean potential generated by all  
4 nucleons. During beta decay, induced by weak interaction, a proton (neutron) in an occupied  
5 orbital decays into a neutron (proton) in a previously unoccupied orbital. During the decay  
6 the nucleon spin may remain the same (Fermi) or may be flipped (Gamow-Teller (GT)). It  
7 is only partially understood why matrix elements of spin-operators, such as the ones for GT  
8 and magnetic transitions and moments, are systematically reduced in the nuclear medium  
9 compared to the single particle values. This renormalisation is often taken into account by  
10 using an effective operator. This is of wider importance since GT beta-decay matrix elements  
11 are relevant in various astrophysical scenarios as well as for double-beta decay[1, 2].

12 In most nuclei measuring the full GT strength is difficult since it is quite fragmented  
13 and only partially accessible in  $\beta$ -decays.  $^{100}\text{Sn}$ , the heaviest doubly magic nucleus with  
14 equal proton and neutron number, is unique since the decay is predicted to dominantly  
15 populate a single state. We measured half-life and Q-value for the GT decay of  $^{100}\text{Sn}$ . We  
16 find the smallest log-ft value of all known beta-decays and a large GT strength establish-  
17 ing it as superallowed GT decay. Large scale shell model calculations that include up to  
18 five particle-hole excitations incorporating the well established renormalisation due to short  
19 range correlations (higher-order effect) allow for a consistent description of the observed  
20 GT-strength for this unique showcase nucleus providing firm and conclusive evidence for the  
21 robust N=Z=50 shell closures at  $^{100}\text{Sn}$ .

## 22 II. INTRODUCTION

23 GT-transitions, in which a proton (neutron) is transformed into a neutron (proton) while  
24 at the same time flipping its spin, are an important spin-isospin degree of freedom in atomic  
25 nuclei. Most of the GT-strength is found in the collective Gamow-Teller Giant Resonance  
26 (GTGR), which is typically a broad structure fragmented over many states. Experiments  
27 investigating the GT-strength via  $\beta$ -decay or charge exchange reactions observed only about  
28 half of the total Ikeda sum-rule [3] in many nuclei. This phenomenon is referred to as *GT-*  
29 *quenching* [4, 5] and has been intensively discussed. The most popular mechanisms for

30 the GT-quenching are particle-hole correlations moving strength to energies at 20-50 MeV,  
 31 substantially above the GTGR, or the coupling to  $\Delta$ -hole excitations (where a nucleon is  
 32 excited to the  $\Delta$ -resonance) moving strength to energies above 300 MeV. Recently new  
 33 (p,n) and (n,p) reaction studies have indeed discovered substantial amounts of GT-strength  
 34 as high as 50 MeV such that about 85-90 % of the sum-rule is exhausted, leaving little room  
 35 for  $\Delta$ -hole excitations [6]. In most calculations that involve truncated model spaces or the  
 36 full major oscillator shell a GT-operator renormalized by a factor  $\approx (1/1.26)$  is used, setting  
 37 artificially the weak coupling constants  $g_A \approx g_V$ , to account for correlations not included in  
 38 the model space. While in charge exchange reactions the full GTGR is accessible,  $\beta$ -decay  
 39 studies can only observe the fraction of the total GT-strength within the decay energy  
 40 window. Near the proton dripline this window becomes larger since the Coulomb repulsion  
 41 shifts the proton levels up. Nevertheless, it is experimentally challenging to detect all small  
 42 GT fragments [6, 7]. The doubly magic  $N = Z$  nucleus  $^{100}\text{Sn}$  provides an experimentally  
 43 and theoretically very clean situation. The closed  $N = 50$  and  $Z = 50$  shells reduce the  
 44 effect of long-range correlations, decreasing the amount of fragmentation of the GTGR. At  
 45 the same time the energy window for beta decay is  $\sim 7.4$  MeV [8] and thus most of the  
 46 GTGR is accessible. In an extreme single particle picture the only possible GT-transition  
 47 is the decay of a proton in a completely filled  $g_{9/2}$  shell to a neutron in an empty  $g_{7/2}$  shell  
 48 since all other final states are Pauli blocked and no levels above  $Z = 50$  are occupied. This  
 49 would lead to a GTGR consisting of only a single  $1^+$  level with a quenched GT-strength  
 50 of  $B_{GT} \approx 10$  taking into account the standard renormalisation factor of 0.75 of the GT  
 51 matrix element originating from short-range correlations [9]. This unique situation is met  
 52 nowhere else in the Segré chart and has been termed super-allowed GT-decay [10]. Even  
 53 in more realistic models, including particle-hole correlations, the GT-decay of the ground  
 54 state of  $^{100}\text{Sn}$  is predicted to populate with more than 95% a single  $1^+$  state in  $^{100}\text{In}$  at an  
 55 excitation energy of about  $3\text{MeV}$ . In these calculations an unquenched  $B_{GT}$  of around 8-14  
 56 [11, 12, 13, 14] is obtained, leading to quenched predictions of 5-7 [12, 14].

57 The production and study of the decay properties of  $^{100}\text{Sn}$  has been the aim of several  
 58 experiments [15, 16, 17, 18, 19], where only a few  $^{100}\text{Sn}$  nuclei were uniquely identified.  
 59 Here we report on a new measurement of the half-life and Q-value from 259 identified  $^{100}\text{Sn}$   
 60 nuclei resulting in the smallest log-ft value of any  $\beta$ -decay on the nuclear chart and a large  
 61 GT-strength establishing the robustness of the  $N=Z=50$  shell closure in  $^{100}\text{Sn}$ .

### 62 III. EXPERIMENTAL DETAILS

63 The experiment was performed at the GSI Helmholtzzentrum Darmstadt, Germany. A  
 64  $^{124}\text{Xe}$  beam of  $E_{kin} = 1\text{AGeV}$  with one second long spills of  $3 \cdot 10^9$  ions every three seconds  
 65 was directed onto a beryllium target placed in front of the fragment separator (FRS) [20].  
 66 Proton-rich nuclei were produced via relativistic projectile fragmentation and guided to the  
 67 final FRS focal plane while being identified event by event (see Figure 1). The correct  
 68 identification was verified by observing the gamma radiation depopulating known isomers,  
 69 e.g. the  $8^+$  isomer in  $^{98}\text{Cd}$ . In total 259  $^{100}\text{Sn}$  nuclei were unambiguously identified, with a  
 70 production rate of 0.75 per hour, yielding a production cross-section of  $5.8(21)\text{pb}$ .

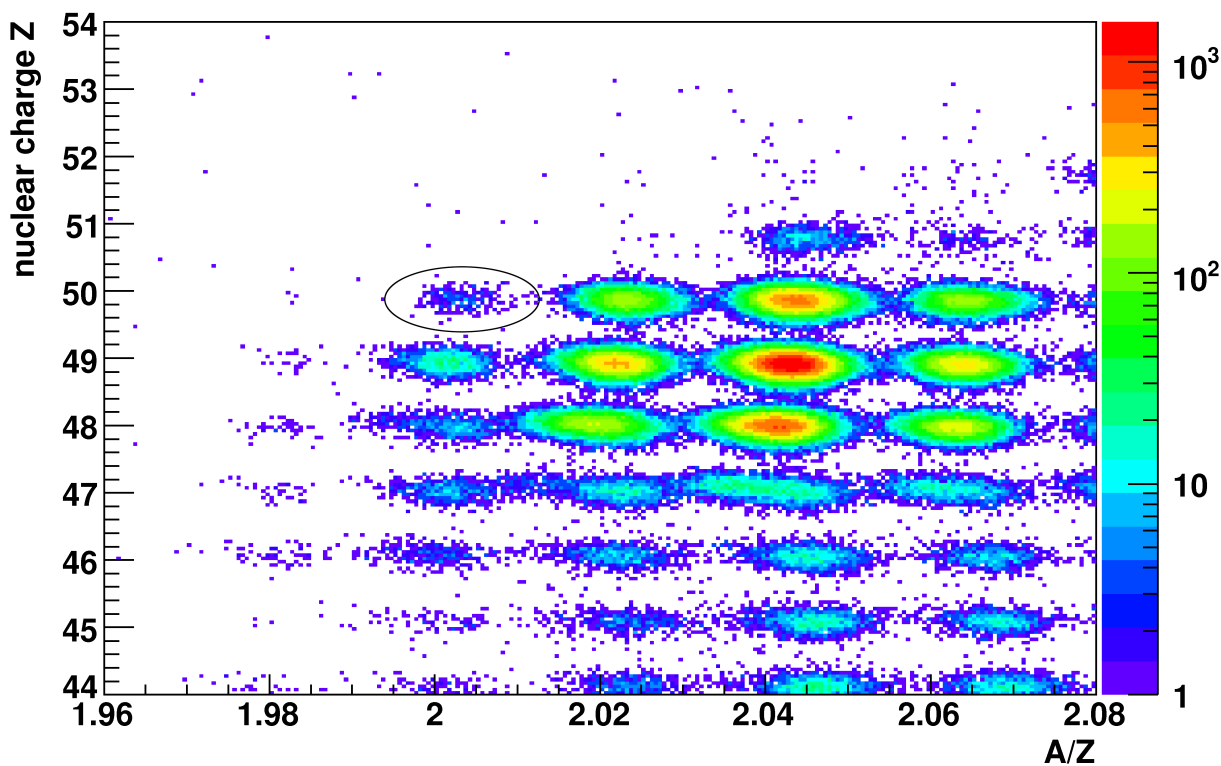


FIG. 1: Particle identification plot with respect to the nuclear charge  $Z$  and the mass-to-charge ratio  $A/Z$  for the full statistics of the  $^{100}\text{Sn}$  fragment separator setting. In total 259  $^{100}\text{Sn}$  nuclei (indicated) have been unambiguously identified. Resolutions (FWHM) in mass of  $\Delta A = 0.32$  and nuclear charge  $\Delta Z = 0.25$  were obtained.

71 The ions were implanted into a stack of highly segmented silicon strip detectors which

72 was surrounded by the RISING array of 105 Germanium detectors in the stopped beam  
73 configuration [21] in order to detect emitted gamma rays with high efficiency.

#### 74 IV. ANALYSIS AND RESULTS

75 Following a  $^{100}\text{Sn}$  implantation in a pixel of the implantation zone all decay events were  
76 recorded that occurred within 15 seconds in this and the directly neighboring pixels in three  
77 dimensions. During this correlation time it was possible to assign 126 decay chains to 163  
78  $^{100}\text{Sn}$  implantations. A maximum likelihood (MLH) analysis with a maximum of three decay  
79 events during the correlation time was used to analyze these decay chains. The half-life of  
80  $^{100}\text{Sn}$  was deduced as  $1.16(20)\text{s}$  in the MLH analysis using literature values for the lifetimes  
81 of the daughter nuclei. The measurement is much more accurate than previous experiments  
82 yielding  $0.94_{-0.27}^{+0.54}\text{s}$  [17] and  $0.55_{-0.31}^{+0.70}\text{s}$  [19]. In Figure 2 the decay curve for  $^{100}\text{Sn}$  is shown.

83 Figure 3 shows the gamma-ray spectrum observed in coincidence with decay events fol-  
84 lowing  $^{100}\text{Sn}$  implantations. The five newly observed transitions are associated with the  
85 depopulation of excited states in the daughter nucleus  $^{100}\text{In}$ .

86 The statistics was only sufficient to establish a coincidence between the 436 keV and 96  
87 keV transitions and thus it is impossible to deduce an unambiguous level scheme of  $^{100}\text{In}$ .  
88 Within uncertainties the transitions could have the same intensity, which would allow for a  
89 single cascade of 5 transitions from the excited  $1^+$  state to the ground state. However, this  
90 would lead to an excitation energy of more than  $4\text{MeV}$ , which is higher than the  $2.5(5)\text{MeV}$   
91 predicted with realistic shell model calculations (see e.g. [22, 23]). The large uncertainties  
92 of the observed intensities also allow for two parallel cascades originating from this  $1^+$  state.

93 Figure 4 shows the relevant level scheme for  $^{100}\text{In}$  obtained from large scale shell model  
94 calculations (LSSM). In this approach  $^{100}\text{Sn}$  is not treated as an inert doubly-magic core  
95 but excitations across the  $Z=N=50$  shell closure were allowed within the  $N=4$  harmonic  
96 oscillator shell with an effective interaction derived from nucleon-nucleon potentials by a  
97 standard many-body approach for a  $^{80}\text{Zr}$  core [24] (see supplementary section). The decay  
98 relevant states of two multiplets are shown,  $(1^+, 2^+, 3^+)$  originating from the  $\pi g_{9/2}^{-1} \otimes \nu g_{7/2}^1$   
99 configuration and  $(2^+, 3^+, 4^+, 5^+, 6^+)$  from the  $\pi g_{9/2}^{-1} \otimes \nu d_{5/2}^1$  configuration. In this calculation  
100 it is assumed that the  $\nu d_{5/2}$  level in  $^{101}\text{Sn}$  lies 172 keV below the  $\nu g_{7/2}$  level [25]. However,  
101 the final conclusions do not change if the reverse situation is assumed, as suggested in [26].

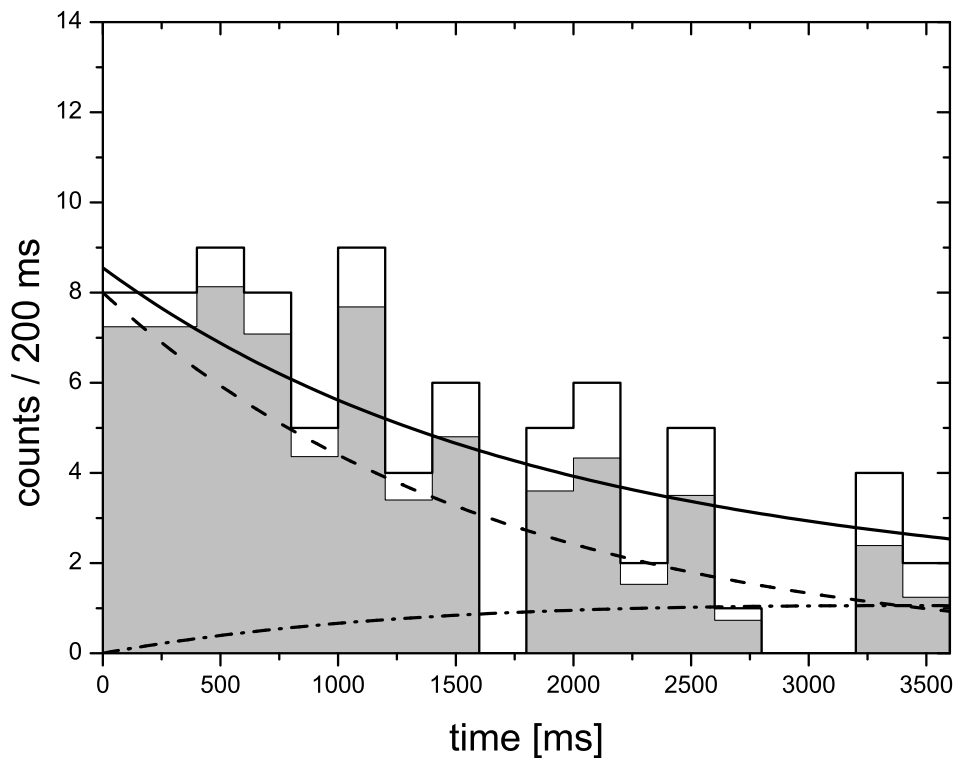


FIG. 2: Decay curve of  $^{100}\text{Sn}$  according to the obtained half life together with the number of  $^{100}\text{Sn}$  decays among all first decay events in 200ms time bins after implantation. Time distribution of first decays after the implantation of  $^{100}\text{Sn}$  nuclei (solid line histogram). Decay curves resulting from the MLH analysis are shown individually for  $^{100}\text{Sn}$  (dashed) and its daughter  $^{100}\text{In}$  (dash-dotted) together with the sum of these decay curves (solid) taking into account a small amount of random background. The light gray filled histogram represents the actual  $^{100}\text{Sn}$  content according to the MLH analysis.

102 The predictions reflect the observed gamma-ray transitions well if the high energy 2048 keV  
 103 transition populates the lowest  $2^+$  state, which decays to lower lying states via the 436, 141,  
 104 and 96 keV transitions ending either in the  $6^+$  ground state or a low lying isomeric state  
 105 with unobserved decay. In this scenario the second  $2^+$  state is populated with the 1297 keV  
 106 transition and decays to the lower lying  $2^+$  and  $3^+$  states. This may lead to a fragmentation  
 107 of the intensities making it impossible to observe these transitions in the present experiment.  
 108 This scenario is supported by the measurement of the total gamma-ray energy in a former

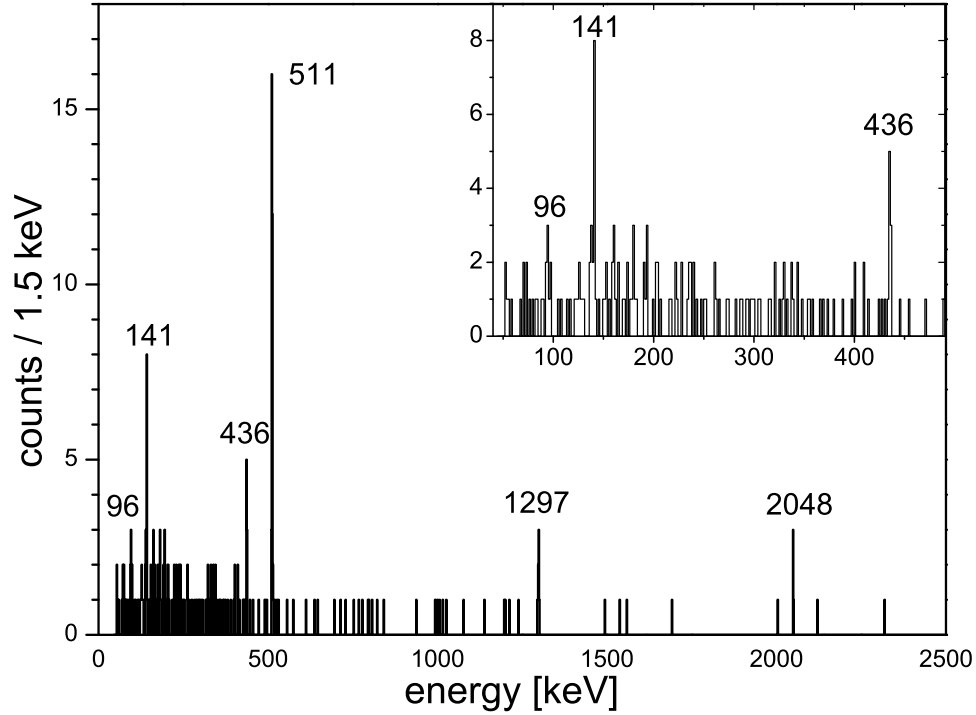


FIG. 3: Energy distribution of the gamma-radiation observed in coincidence with the beta decay of  $^{100}\text{Sn}$ . The spectrum contains with a probability of 65% radiation directly following  $^{100}\text{Sn}$  decay events. The other contributions are uncorrelated background decays and daughter decays of  $^{100}\text{In}$ . None of the observed lines corresponds to known transitions from these minor contributions. The measured absolute intensities of the five lines with the energies 96, 141, 436, 1297 and 2048 keV are  $79 \pm 40$ ,  $100 \pm 31$ ,  $59 \pm 22$ ,  $72 \pm 26$  and  $53 \pm 26$  corrected for the emission of conversion electrons assuming M1-transitions and  $141 \pm 70$ ,  $122 \pm 37$ ,  $59 \pm 22$ ,  $72 \pm 26$  and  $53 \pm 26$  assuming E2-transitions, respectively.

109 experiment with a BGO-detector [17], the known mass difference of  $^{100}\text{Sn}$  and  $^{100}\text{In}$  [16]  
 110 and our observation of a single event of  $\beta$ -delayed proton emission and is fully consistent  
 111 with the expectation that dominantly a single  $1^+$  state is populated in the decay. Further  
 112 details are given in the supplementary section.

113 As a key feature of this experiment, the kinetic energy of the decay positrons fully ab-  
 114 sorbed in the compact silicon detector array was measured. The spectrum resulting from the

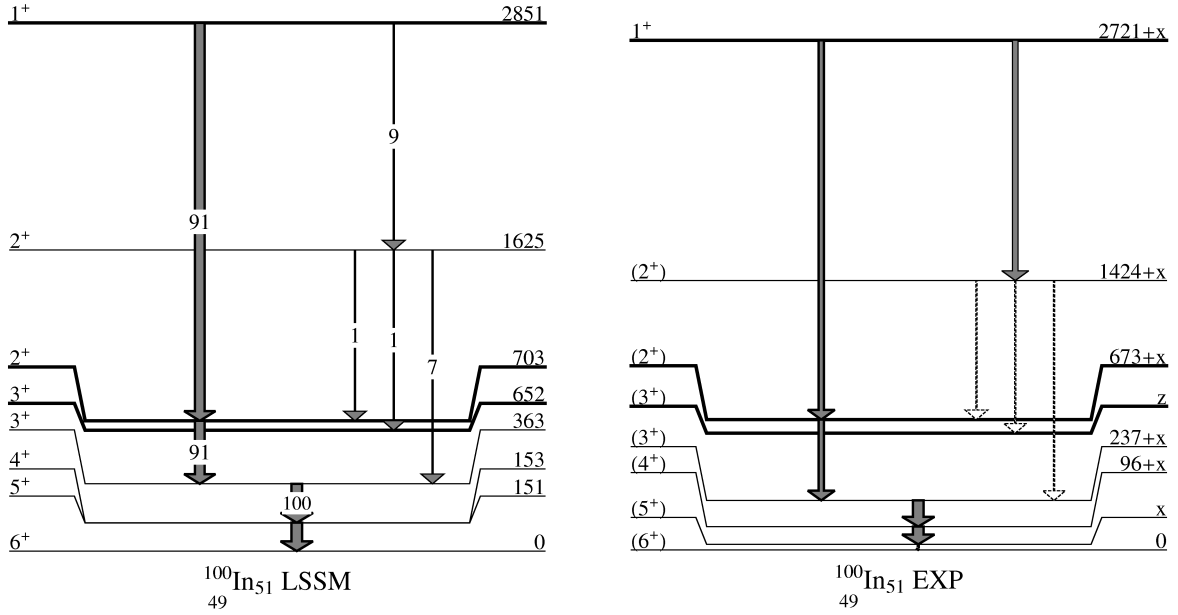


FIG. 4: *Left:* Large scale shell model (LSSM) calculation of the low lying excited states in  $^{100}\text{In}$ . The energy of the states is given in keV and the quantum numbers denoting the states are spin and parity. Populated levels with almost pure  $\pi g_{9/2}^{-1} \otimes \nu g_{7/2}^1$  configuration are indicated with bold lines while the remaining levels are part of the  $\pi g_{9/2}^{-1} \otimes \nu d_{5/2}^1$  multiplet. Explicit numbers for the intensities of selected transitions relative to the strongest transition are shown. *Right:* Possible level scheme resulting from the most likely interpretation of the observed five gamma ray transitions in  $^{100}\text{In}$ . Since at least one low energy transition might not have been observed in the cascade on the left if using only M1-transitions the energy of the levels might have a systematic shift up to  $x = 80\text{keV}$  due to the experimental detection threshold. The dashed transitions are most likely below the detection limit of the experiment.

115 summed up energies of the uninterrupted  $\beta$ -traces up to three seconds after a  $^{100}\text{Sn}$  implan-  
 116 tation is shown in Figure 5. It was fitted using a MLH analysis based on a single component  
 117 beta decay phase space function to determine the endpoint energy in the decay of  $^{100}\text{Sn}$ . For  
 118 the fit of the end point energy only data in the energy region between  $400\text{keV}$  and  $2600\text{keV}$   
 119 were used. In the analysis corrections were applied for the occurrence of converted low en-  
 120 ergy transitions emitted during the deexcitation of the daughter nucleus  $^{100}\text{In}$ , for emitted  
 121 Bremsstrahlung of the positrons, and for the annihilation of positrons in flight before the  
 122 deposition of their total kinetic energy. The measured value of the endpoint energy of the  
 123  $\beta$ -decay populating a single final state in the daughter nucleus  $^{100}\text{In}$  was determined to be



124  $3.29(20)MeV$ . The corresponding fraction of electron capture decays was calculated to be  
 125 13% of all  $^{100}Sn$  decays. Assuming the four lowest  $1^+$  states in  $^{100}In$  with the energy splitting  
 126 and the relative GT-strength from the LSSM calculation (see supplementary section), the  
 127 measured beta-spectrum would yield an endpoint energy for the transition to the lowest  $1^+$   
 128 state that is larger by  $73keV$ .

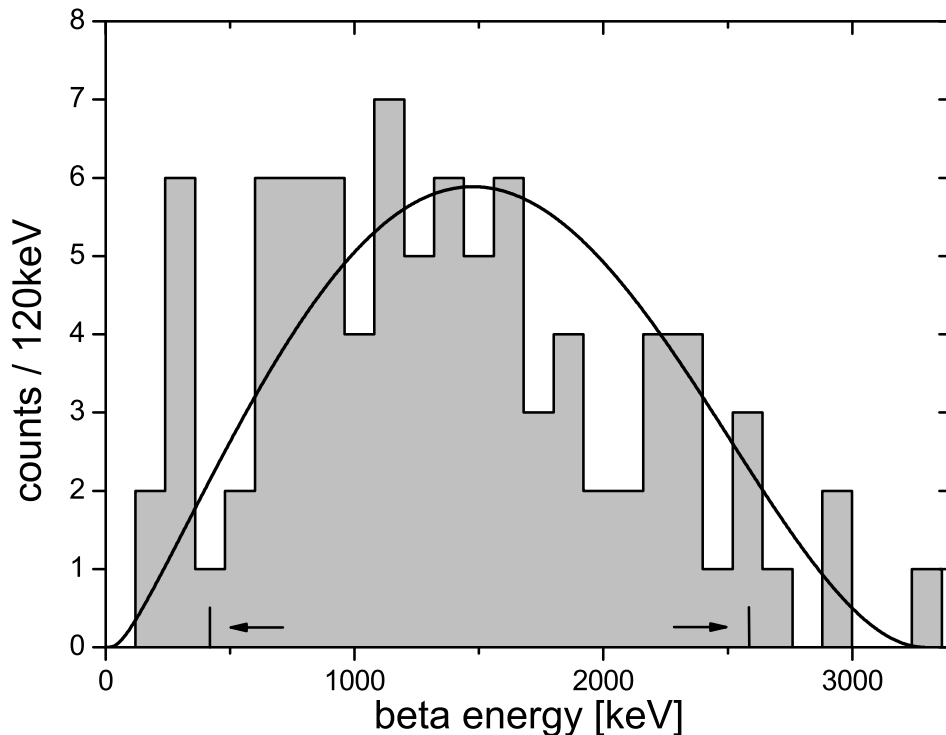


FIG. 5: Distribution of the positron energies emitted in the beta decay of  $^{100}Sn$ . The spectrum contains only decay events which can be assigned to  $^{100}Sn$  decays with a probability of at least 75%. The MLH fit was applied to the region between 400keV and 2600keV which is indicated with markers. The solid curve illustrates the shape of the best fitting single component beta decay phase space function determined in a MLH analysis.

## 129 V. DISCUSSION

130 Using the measured half-life and the end point energy a log-ft value of  $2.62^{+0.13}_{-0.19}$  was  
 131 extracted, which is the smallest known log-ft value for any beta decay in the nuclear chart.

132 Thus, the Gamov-Teller decay of  $^{100}\text{Sn}$  has much larger strength than the known  $0^+ \rightarrow 0^+$   
133 superallowed Fermi decays of  $N = Z$  nuclei and is indeed a *superallowed Gamow-Teller decay*  
134 [27]. The extracted Gamow-Teller strength  $B_{GT}$  of the  $^{100}\text{Sn}$  ground state decay to the single  
135 excited  $1^+$  state in  $^{100}\text{In}$  is  $B_{GT} = 9.1_{-2.3}^{+4.8}$ . This is surprisingly large but within uncertainties  
136 still consistent with the value of  $B_{GT} = 5.8_{-3.2}^{+5.5}$  deduced from previous results of  $Q_{EC}$  and  
137  $T_{1/2}$  [28]. The uncertainty of the new  $B_{GT}$  value is dominated by the uncertainty of the beta  
138 endpoint energy. The extraction of the strength was performed under the assumption of  
139 the GT-decay into only one final  $1^+$  state in  $^{100}\text{In}$ . However, if other  $1^+$  states at excitation  
140 energies above the observed state would also be populated, the summed GT-strength would  
141 only increase while the  $B_{GT}$  for the first excited  $1^+$  state would decrease. It would have  
142 been difficult to observe such states since the reduced phase space for higher lying  $1^+$  states  
143 would have lead to a strongly reduced population.

144 The LSSM calculations which take into account long-range correlations across the  
145  $N=Z=50$  doubly magic shell closure and include up to five particle-hole excitations (see  
146 supplementary section for details) yield a total summed GT-strength of all possible final  
147 states in the daughter nucleus up to  $60\text{MeV}$  of  $B_{GT} = 8.19$  applying the standard renor-  
148 malisation due to short-range correlations.

149 This quenching of the GT-strength takes into account those correlations not included in  
150 the model space, in particular higher energy particle-hole excitations or excitation of the  
151  $\Delta$  resonance. Higher-order correlations from beyond the  $0 \hbar\omega$  valence space are well docu-  
152 mented in the  $N=1,2,3$   $p, sd, pf$  harmonic oscillator shells [9] amounting to a reduction of  
153 the GT matrix element by a factor of about 0.75 in medium-heavy nuclei. This is a property  
154 of the wave functions and in line with evidence for the M1 operator and spectroscopic factors  
155 (see Ref.[9] for a summary). Their absence or reduction in the  $N=4$   $gds$  shell therefore is  
156 unlikely.

157 A quenched GT-strength of  $B_{GT} = 7.82$  is predicted in the experimental  $Q_{EC}$  window  
158 of  $7.03(20)$  MeV. This corresponds to a reduction of the total renormalised GT strength  
159 of the extreme single particle estimate of  $B_{GT} \approx 10$  by 18% for excitation energies up to  
160  $60\text{MeV}$  and by 22% in the  $Q_{EC}$  window. It is due to mixing in the  $gds$  harmonic oscillator  
161 shell, i.e. emptying of the proton  $g_{9/2}$  orbital and pre-filling of the neutron  $g_{7/2}$  orbital.  
162 The occupation numbers of the two orbitals which are linked by the GT operator of course  
163 directly influence the strength of the transition matrix element.

164 However, the prediction is that the largest fraction of the strength remains located in the  
 165 first excited  $1^+$  state which agrees with earlier calculations [11, 27]. Nevertheless, according  
 166 to the LSSM calculations performed in this work it is reasonable to consider that several  $1^+$   
 167 states in  $^{100}\text{In}$  are populated in the decay of  $^{100}\text{Sn}$ . If we take from the LSSM calculation  
 168 the four lowest  $1^+$  states in  $^{100}\text{In}$  with their energy splitting and relative GT-strength (see  
 169 supplementary section), the value of  $B_{GT}(1_1^+) = 9.1_{-2.3}^{+4.8}$  (assuming a single  $1^+$  state) would  
 170 be reduced to  $B_{GT}(1_1^+) = 7.6_{-2.1}^{+2.6}$  for the first excited  $1^+$  state using the experimental log-ft  
 171 value. The corresponding summed GT-strength would be  $\sum_{i=1}^4 B_{GT}(1_i^+) = 9.9_{-2.8}^{+3.4}$  using the  
 172 experimental log-ft value and the theoretical distribution of strength.

173 The quenched LSSM result of  $B_{GT} = 5.68$  of the GT-strength for this first excited  $1^+$  state  
 174 agrees well with the value  $B_{GT} = 7.6_{-2.1}^{+2.6}$  within the statistical uncertainty extracted from  
 175 the experimental log-ft value under above assumptions. The experimental concentration of  
 176 most of the GT strength in the first excited  $1^+$  state with marginal fragmentation clearly  
 177 classifies the  $^{100}\text{Sn}$  GT decay as super-allowed. The calculation shows that the large GT-  
 178 strength of the first excited  $1^+$  state coincides with relatively pure wave functions of the  
 179  $^{100}\text{Sn}$  ground state and of the first excited  $1^+$  state in  $^{100}\text{In}$  consisting significantly of the  
 180  $g_{9/2}^{10} \otimes g_{7/2}^0$  ( $xxx\%$ ) and the  $g_{9/2}^9 \otimes g_{7/2}^1$  ( $yyy\%$ ) component, respectively. The high purity of  
 181 the wave functions establishes for the first time the simultaneous robustness of the  $Z=50$   
 182 and  $N=50$  shell closures in  $^{100}\text{Sn}$  only  $\sim 3\text{MeV}$  from the proton dripline corroborating for  
 183  $N=50$  the results of Refs.[29, 30] and clearly excludes substantial coupling to the continuum  
 184 of unbound states.

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## 222 A. Experimental details

## 223 1. Fragment separator (FRS)

224 Exotic nuclei were produced via relativistic projectile fragmentation. A  $^{124}\text{Xe}$ -beam with  
 225 an energy of  $1\text{AGeV}$  per nucleon was directed onto a  $4008\text{ mg/cm}^2$  beryllium target. A pro-  
 226 duction cross section of  $5.8(21)\text{pb}$  [S1] for  $^{100}\text{Sn}$  was determined to be compared with  $11(4)\text{pb}$   
 227 from [S2]. The FRS [S3] was used to separate the various reaction products. It was operated  
 228 with an achromatic focus at the final focal plane and a wedge shaped aluminium degrader  
 229 in the dispersive midplane (F2) was inserted to achieve additionally a spatial separation  
 230 with respect to the mass-to-charge ratio  $A/Q$  and the nuclear charge  $Z$  of identified ions.  
 231 In the focal plane F1 after the first quarter of the FRS an additional  $2\text{g/cm}^2$  Al degrader  
 232 was placed to keep the countrate at F2 at a tolerable level. Scintillators and tracking ion-  
 233 ization chambers at the dispersive midplane (F2) and two multiwire proportional counters  
 234 and two time-projection chambers at F4 were used to determine the fragments magnetic  
 235 rigidity  $B\rho = (A/Q)m\gamma v$  by measuring its horizontal positions and angles. Multiple time of  
 236 flight measurements with scintillators at F2 and F4 yielded the velocity  $v$  and the Lorentz  
 237 factor  $\gamma$  and thus the fragments  $A/Q$ . The total time of flight was approximately  $200\text{ns}$   
 238 in the fragments rest-frame. At the relativistic energies used here the fragments are com-  
 239 pletely ionized ( $Q = Ze$ ). An independent determination of the nuclear charge by energy  
 240 loss measurements in two multiply sampling ionization chambers (MUSIC) at F4 led to an  
 241 unambiguous identification of the nuclide. Resolutions (FWHM) in mass of  $\Delta A = 0.32$  and  
 242 nuclear charge  $\Delta Z = 0.25$  were obtained. At the final focal plane of the FRS the ions were  
 243 slowed down in a degrader of adjustable thickness and stopped in an implantation detector.

## 244 2. Implantation detector and RISING gamma detection array

245 The compact  $4\pi$  silicon implantation and beta absorber detector (SIMBA) is sensitive  
 246 to  $\beta$ -,  $\alpha$ -, and proton-decay of the heavy ions [S4]. The implantation zone at the center of  
 247 SIMBA consists of 3 highly segmented  $0.7\text{mm}$  thick double sided silicon strip detectors with  
 248 a total number of 1920 pixels per  $60\text{mm} \times 40\text{mm}$  detector. This high granularity enables

249 correlating the implantation and successive decays of a nucleus for times up to several  
 250 seconds. In front of and behind the implantation zone stacks of ten  $1mm$  thick silicon  
 251 detectors with lower segmentation were mounted as a total energy calorimeter, stopping  
 252 beta particles with a maximum energy of  $5MeV$  and yielding an average energy resolution  
 253 of  $100keV$  FWHM due to summing up of the energies for an ordinary  $\beta$  decay depositing  
 254 energy in about 6 silicon layers of SIMBA. The surrounding RISING array for the detection  
 255 of emitted gamma radiation has an energy resolution of about  $3.0keV$  FWHM at  $1332keV$   
 256 and the absolute photo peak efficiency of the whole setup was about 12% at  $662keV$  for  
 257 gamma radiation originating from the implantation zone.

## 258 **B. Analysis details**

### 259 *1. Maximum likelihood (MLH) analysis of decay events*

260 Due to an average efficiency of only 60% for correlating implantations with succeeding  
 261  $\beta$ -decays and due to the presence of uncorrelated background decays from previous implan-  
 262 tations some ambiguity exists in the assignment of each decay event to mother-, daughter-,  
 263 granddaughter-, and background-decays. All possible scenarios with a maximum number of  
 264 three decay events during the correlation time have been incorporated in the MLH analysis.  
 265 The unknown half life of  $^{100}Sn$  was then varied in order to maximize the probability of  
 266 observing the data sample as it was recorded in the experiment. Considering the probability  
 267 distribution for each decay event of the entire data sample a total number of  $83_{-4}^{+2}$  of all  
 268 observed decays can be attributed to decays of  $^{100}Sn$ .

### 269 *2. Decay gamma ray spectrum and level scheme of $^{100}In$*

270 Figure 3 shows the gamma-ray spectrum observed in coincidence with decay events which  
 271 originate with a probability of 65% from excited states in  $^{100}In$ . A 35% contribution origi-  
 272 nates from daughter decays or uncorrelated background. The strongest transition from the  
 273 daughter decay with  $1004keV$  is barely visible in the spectrum. Furthermore none of the  
 274 visible gamma-rays corresponds to lines in a high statistics spectrum obtained from uncor-  
 275 related background decays. The observed intensities are within the very limited statistics  
 276 all consistent with  $\sim 70$   $^{100}Sn$  decays corresponding to a correlation window of 4 seconds.

277 The proposed scenario of the depopulation of excited states in  $^{100}\text{In}$  as shown in Figure 4  
 278 is consistent with the observation of a total gamma-ray energy of 2.76(43) MeV measured  
 279 with a total absorption BGO-detector [S2] in a former experiment. In addition, the mea-  
 280 sured mass difference of  $^{100}\text{Sn}$  and  $^{100}\text{In}$  [S5] combined with our measured  $Q_{EC}$  value of 3.29  
 281 MeV yields an excitation energy of the  $1^+$  state of 2.6(10) MeV. A third indication is the  
 282 observation of one  $^{100}\text{Sn}$  decay event, where 1.49 MeV has been deposited in a single pixel  
 283 in coincidence with a  $\beta$  particle track. Since a  $\beta$ -particle cannot deposit that much energy  
 284 we attribute this event to a  $\beta$ -delayed proton emission. Subtraction of the average energy  
 285 deposited by a  $\beta$ -particle in a single pixel of 0.17(14) MeV yields a proton energy of 1.32(14)  
 286 MeV. The natural assumption is that the  $1^+$  state, fed by  $\beta$ -decay, is depopulated to a  
 287 fraction  $< 1\%$  by a proton decay to the  $^{99}\text{Cd}$  ground state. By adding the proton separation  
 288 energy of  $^{100}\text{In}$   $S_p = 1.61(31)\text{MeV}$  [S6] we arrive at an excitation energy of the  $1^+$  state in  
 289  $^{100}\text{In}$  of 2.93(34) MeV. For the partial proton-decay half-life we estimate [S7] 40ps with more  
 290 than a factor of 10 uncertainty for an L=2 transition to the  $5/2^+$  ground state of  $^{99}\text{Cd}$ . This  
 291 is about a factor  $10^4$  larger than for a 2048 keV  $M1$ -transition with 1 Weisskopf unit. Thus  
 292 a proton decay branch in the range of 0.1% to 1.0% is not unexpected.

### 293 3. Determination of the beta endpoint energy

294 For all decay events which could be assigned to a  $^{100}\text{Sn}$  decay with a probability of at least  
 295 75% the uninterrupted traces of the decay positrons in the implantation detector were recon-  
 296 structed and checked for integrity, i.e. for possible positron escapes with incomplete energy  
 297 deposition and interference of the trace with single Compton events of gamma rays emitted  
 298 in coincidence. The beta spectrum resulting from the summed up energies shown in Figure  
 299 5 was fitted using a MLH analysis based on a single component beta decay phase space  
 300 function to determine the beta endpoint energy in the decay of  $^{100}\text{Sn}$ . The region below  
 301 400keV was neglected to avoid the narrow structure (7 counts) around 250keV, whose origin  
 302 is only partially understood. Two decays within this region which occurred without a coin-  
 303 cident observation of any gamma radiation might be attributed to uncorrelated background  
 304 noise rather than real decay events since these low energies are not far from the trigger  
 305 threshold. Another two events might come from electron capture decays of  $^{100}\text{Sn}$  (fraction  
 306 13%) accompanied by an energy deposition in the spatial three dimensional correlation win-



307 dow originating from Compton scattered gamma rays emitted during the depopulation of  
 308 excited states in the daughter nucleus  $^{100}\text{In}$ . Thus the narrow structure around  $250\text{keV}$  has  
 309 a large systematic uncertainty and was not considered in the fitting procedure. The MLH  
 310 fit of the end-point energy would be strongly influenced by the highest energy events, since  
 311 obviously the end-point of a one component phase space would have to be higher than the  
 312 highest energy events. However, one cannot exclude that the few events with energies above  
 313  $2600\text{keV}$  could originate from the daughter (probability 13%) and uncorrelated background  
 314 decays (probability 12%). Therefore these high energy events were also excluded from the  
 315 MLH fit. However, the choice of the upper limit for the fit in the tail of the distribution as  
 316 indicated in Figure 5 has only a weak influence on the result which is much less than the  
 317 statistical uncertainty.  
 318 The corresponding fraction of electron capture decays was calculated from the determined  
 319 beta endpoint energy to be 13% of all  $^{100}\text{Sn}$  decays [S8].

### 320 C. Interpretation of the data

#### 321 1. Compilation of various predictions for the GT-strength in the decay of $^{100}\text{Sn}$

322 The basic estimate for the GT-strength in the decay of  $^{100}\text{Sn}$  comes from the extreme  
 323 single particle shell model (ESPM) where no correlations between the nucleons are taken  
 324 into account. The GT transition strength is given by  $B_{GT}^{ESM} = \frac{4\ell}{2\ell+1} \cdot (1 - \frac{N_{\nu g7/2}}{8}) \cdot N_{\pi g9/2}$  [S9].  
 325  $\ell = 4$  denotes the involved orbital angular momentum.  $N_{\pi g9/2} = 10$  and  $N_{\nu g7/2} = 0$  are the  
 326 occupation numbers of the initial proton orbital and the final neutron orbital, respectively.  
 327 The extreme single particle estimate increases linearly with the orbital angular momentum  
 328  $\ell$ , and takes its maximum value for a given  $\ell$  for a filled  $\ell + 1/2$  and empty  $\ell - 1/2$  orbit.  
 329 The constraints for observing the maximum value in  $\beta^+/\text{EC}$  decay are given by the line of  
 330 stability and the proton dripline which exclude doubly magic nuclei below and above  $N=Z=50$   
 331 from observation making  $^{100}\text{Sn}$  a unique case in the Segré chart.

332 The experimental  $B_{GT}$  value is substantially lower than the pure single particle estimate of  
 333  $B_{GT,sp} = 17.78$ , which assumes the decay of one of the  $N_{\pi} = 10$  protons occupying the full  
 334  $g_{9/2}$  orbital into an empty  $g_{7/2}$  neutron orbital with  $N_{\nu} = 0$ . This is not unexpected since  
 335 particle-hole correlations within the full  $gds$  shell as well as short range correlations shifting

336 strength to even higher energies would lead to a reduced GT-strength.

337 Table I summarizes various theoretical results for the prediction of the GT strength in the  
 338 decay of  $^{100}\text{Sn}$ . The quenched values are obtained by multiplying the square of the quenching  
 339 factor to the Gamow-Teller strength.

Model	Ref	unquenched	quenched	quenching factor
ESPM	[S9]	17.78	10.00	0.75
MCSM	[S10]	10.3	6.5	0.79
QRPA	[S11]	8.95		
FFS	[S11]	7.63		
extrapol.	[S12]	9.8	5.2	0.75
SM+corr.	[S13]	14.2		
LSSM	this work	$\sim 13.90$	$\sim 7.82$	0.75
LSSM (only $1_1^+$ )	this work	10.10	5.68	0.75

TABLE I: Predicted  $B_{GT}$  values of the strength accessible in the  $Q_{EC}$  energy window in the decay of  $^{100}\text{Sn}$ . The measured  $B_{GT}$  is  $9.1_{-2.3}^{+4.8}$  assuming a single final state dominantly populated in the decay of  $^{100}\text{Sn}$ .

## 340 2. Details of shell model calculations

341 The large scale shell model (LSSM) calculations in this work for the determination of the  
 342  $^{100}\text{In}$  level scheme and for the determination of the GT-strength distribution in the decay of  
 343  $^{100}\text{Sn}$  depending on possible final states in the daughter nucleus  $^{100}\text{In}$  were performed in the  
 344 proton-neutron  $\pi\nu(g, d, s)$  model space with a  $^{80}\text{Zr}$  core making use of the Strasbourg codes  
 345 ANTOINE and NATHAN [S14, S15]. Up to 5 particle - 5 hole excitations ( $5p5h$ ), truncation  
 346 level  $t = 5$ , across the Z,N=50 closed shells were allowed. The interaction was derived from  
 347 the CD-Bonn nucleon-nucleon ( $NN$ ) potential [S16], starting from a G-matrix including  
 348 folded diagrams and higher order in-medium corrections for this model space according to  
 349 the prescription of Ref. [S17]. Starting from a global prediction for single particle energies

350 for a hypothetical (spherical)  $^{80}\text{Zr}$  core [S18] monopole corrections were applied to reproduce  
 351 extrapolated single particle/hole energies for  $^{100}\text{Sn}$  [S19] including a shell gap of 6.50 MeV  
 352 for both protons and neutrons. For details see Ref. [S20] where this approach was verified for  
 353 core excited states in  $^{98}\text{Cd}$ . Recently additional support was obtained from the description  
 354 of a core excited isomer in  $^{96}\text{Ag}$  [S21].

355 The resulting  $^{100}\text{In}$  level scheme relevant for the decay pattern of the  $1^+$  state is shown  
 356 in Figure 4 of the main text. E2 polarisation charges  $\delta e = 0.5e$  and effective spin  $g$ -factors  
 357  $g_s^{eff} = 0.75g_s^{free}$  and orbital corrections  $\delta g_l = \pm 0.1$  for protons and neutrons were used for  
 358 M1 transitions. The  $\gamma$  decay pattern of the first  $1^+$  state in this model space as calculated  
 359 from transition strengths excludes E2 transitions in competition to M1 and may be used as  
 360 guideline to arrange the experimentally observed  $\gamma$  rays in a tentative level scheme. The  
 361 disparate branching from the first  $1^+$  state of dominant configuration  $\pi g_{9/2}^{-1} \nu g_{7/2}$  to the  $2_{1,2}^+$   
 362 states in the LSSM calculations may be due to the neglect of  $l$ -forbidden M1 transitions as  
 363 the daughter states have different partitions of the  $\pi g_{9/2}^{-1} \nu g_{7/2}$  and  $\pi g_{9/2}^{-1} \nu d_{5/2}$  configurations.

364 The quenched GT distribution, as shown in Figure 1 of the supplementary section, is  
 365 used to correct the total GT strength for the part observable in the  $Q_{EC}$  window. Note the  
 366 multiple  $1^+$  states above the dominantly populated  $1_1^+$ . The states  $1_{2-4}^+$  between 3.5 and  
 367 4.0 MeV carry 21% of the total strength while the  $1_1^+$  receives 69% of the total strength.  
 368 Nevertheless due to the reduced phase space factor the former is not accessible to the present  
 369 experiment.

370 The total quenched GT strength in an energy window of up to 60 MeV for excited states in  
 371 the daughter nucleus yields a value of 8.19. With a quenched GT-strength of 7.82, in the  
 372  $Q_{EC}$  window of 7.03(20) MeV most of the strength is concentrated. The value determined  
 373 for the first excited  $1^+$  state is 5.68.

### 374 3. Interpretation of the measured GT-strength $B_{GT}$

375 According to the large scale shell model (LSSM) calculations it is reasonable to consider  
 376 that a few  $1^+$  states in  $^{100}\text{In}$  are populated in the decay of  $^{100}\text{Sn}$  with the corresponding  
 377 distribution of the Gamow-Teller strength  $B_{GT}$  as shown in Figure 1 of the supplementary  
 378 section. If we assume from the LSSM calculation the four lowest  $1^+$  states in  $^{100}\text{In}$  with  
 379 the same energy splitting and the same relative GT-strength, the measured beta-spectrum

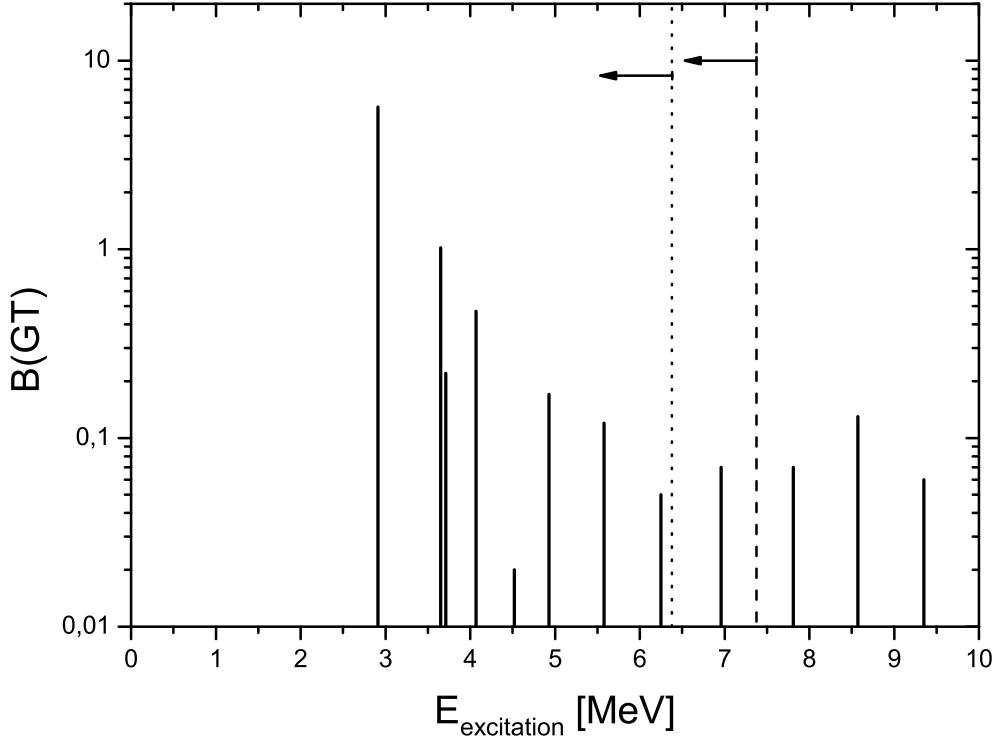


FIG. 1: LSSM calculation of the distribution of the quenched GT-strength in the decay of  $^{100}\text{Sn}$  depending on the excitation energy of possible final  $1^+$  states in the daughter nucleus  $^{100}\text{In}$ . The standard renormalisation factor of 0.75 of the GT-operator was applied to account for short-range correlations. The dashed line indicates the energy window accessible in EC-decays and the dotted line shows the energy window accessible in  $\beta^+$  decays, respectively.

380 would yield an endpoint energy for the transition to the lowest  $1^+$  state that is larger by  
 381  $73\text{keV}$ . This is due to the shift of the centroid of the  $\beta$ -spectrum to lower energies induced  
 382 by the minor contribution of the feeding of the higher lying  $1^+$  states in contrast to the  
 383 assumption that only a single final state is populated.

384 The corrected value of the beta-endpoint energy together with the corrected partial half  
 385 life would result in a  $B_{GT}(1_1^+) = 7.6_{-2.1}^{+2.6}$ , a factor of 1.3 larger than the result of the LSSM  
 386 calculation of 5.68, which is still within the uncertainties. The summed GT-strength is  
 387  $\sum_{i=1}^4 B_{GT}(1_i^+) = 9.9_{-2.8}^{+3.4}$  in comparison to the LSSM value of 7.40.

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